

Scattering

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1 Polarization

Let's describe the electric field \vec{E} in terms of two orthogonal polarization vectors \hat{e}_1 and \hat{e}_2 . Then:

$$\vec{E} = \vec{E} \cdot \hat{e}_1 + \vec{E} \cdot \hat{e}_2$$

and

$$|\vec{E}|^2 = (\vec{E} \cdot \hat{e}_1)^2 + (\vec{E} \cdot \hat{e}_2)^2$$

Thus the power radiated, which is proportional to $|\vec{E}|^2$, is the sum of the powers radiated into the two orthogonal polarizations.

Now in the non-relativistic case, or with \vec{a} parallel to $\vec{\beta}$, $\vec{E} \propto \hat{n} \times (\hat{n} \times \vec{a})$ and so

$$\begin{aligned} \vec{E} \cdot \hat{e}_1 &\propto (\hat{n} \times (\hat{n} \times \vec{a})) \cdot \hat{e}_1 \\ &= (\hat{n} (\hat{n} \cdot \vec{a}) - \vec{a}) \cdot \hat{e}_1 \\ &= (\hat{n} \cdot \vec{a}) \hat{n} \cdot \hat{e}_1 - \vec{a} \cdot \hat{e}_1 \end{aligned}$$

But since \hat{n} is perpendicular to \hat{e}_1 , then

$$\vec{E} \cdot \hat{e}_1 \propto \vec{a} \cdot \hat{e}_1$$

Generally this is the easiest way to compute the power.

2 Thomson scattering

Let the incident wave be described by:

$$\vec{E}(\vec{x}, t) = \hat{e}_0 E_0 \exp\left(i \left[\vec{k} \cdot \vec{x} - \omega t \right]\right)$$

where \hat{e}_0 is a vector describing the polarization. The electric field is incident on an electron and gives it an acceleration:

$$\vec{a} = \frac{-e}{m} \hat{e}_0 E_0 \exp\left(i \left[\vec{k} \cdot \vec{x} - \omega t \right]\right)$$

and then the power radiated into polarization i is:

$$\begin{aligned}\frac{dP_i}{d} &= \frac{e^2}{4\pi c^3} |\vec{a} \cdot \hat{e}_1|^2 \\ &= \frac{e^2}{4\pi c^3} \left(\frac{e}{m} E_0\right)^2 \left(\hat{\epsilon}_0 \exp\left(i\left[\vec{k} \cdot \vec{x} - \omega t\right]\right) \cdot \hat{e}_1^*\right)^2\end{aligned}$$

and the time average is

$$\left\langle \frac{dP_i}{d} \right\rangle = \frac{e^4}{4\pi m^2 c^3} \frac{E_0^2}{2} (\hat{\epsilon}_0 \cdot \hat{e}_i^*)^2$$

Note: by using the complex conjugate of the polarization vector to compute the absolute value, we can include circular polarizations in our results.

The differential scattering cross section is defined as:

$$\frac{d\sigma}{d} \equiv \frac{\text{energy radiated/unit time/unit solid angle}}{\text{incident flux}}$$

and thus

$$\begin{aligned}\frac{d\sigma_i}{d} &= \frac{\frac{e^4}{8\pi m^2 c^3} E_0^2 (\hat{\epsilon}_0 \cdot \hat{e}_i^*)^2}{c E_0^2 / 8\pi} \\ &= \frac{e^4}{m^2 c^4} (\hat{\epsilon}_0 \cdot \hat{e}_i^*)^2 \\ &= r_0^2 (\hat{\epsilon}_0 \cdot \hat{e}_i^*)^2\end{aligned}$$

where

$$r_0 = \frac{e^2}{mc^2}$$

is the classical electron radius.

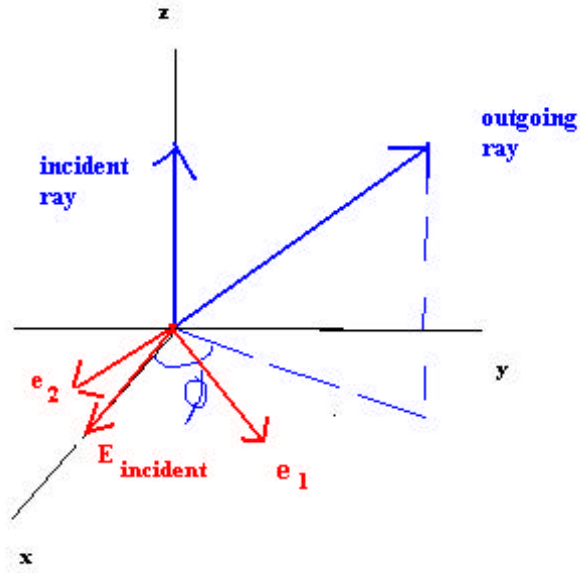
Let's assume the incident wave is incident along the z -axis and is linearly polarized along the \hat{x} -axis: $\hat{\epsilon}_0 = \hat{x}$. An outgoing wave scattered at angle θ may be described in terms of the polarization vectors

$$\hat{e}_1 = \cos\theta (\hat{x} \cos\phi + \hat{y} \sin\phi) - \hat{z} \sin\theta \quad (1)$$

and

$$\hat{e}_2 = -\hat{x} \sin\phi + \hat{y} \cos\phi$$

as shown in the diagram:



Then

$$\hat{\epsilon}_0 \cdot \hat{\epsilon}_1 = \cos \theta \cos \phi$$

and

$$\hat{\epsilon}_0 \cdot \hat{\epsilon}_2 = -\sin \phi$$

Thus:

$$\begin{aligned} \frac{d\sigma}{d} &= \frac{d\sigma_1}{d} + \frac{d\sigma_2}{d} \\ &= r_0^2 (\cos^2 \theta \cos^2 \phi + \sin^2 \phi) \end{aligned}$$

For the perpendicular incident polarization $\hat{\epsilon}_0 = \hat{y}$, we have:

$$\frac{d\sigma}{d} = r_0^2 (\cos^2 \theta \sin^2 \phi + \cos^2 \phi)$$

and thus the differential scattering cross section for unpolarized incident radiation is:

$$\frac{d\sigma}{d} = \frac{r_0^2}{2} (1 + \cos^2 \theta)$$

and thus the total scattering cross section is:

$$\begin{aligned} \sigma &= \int \frac{d\sigma}{d} d = 2\pi \frac{r_0^2}{2} \int_{-1}^{+1} (1 + \mu^2) d\mu = \pi r_0^2 \left(\mu + \frac{\mu^3}{3} \right) \Big|_{-1}^{+1} \\ &= \frac{8\pi}{3} r_0^2 = 0.665 \times 10^{-24} \text{ cm}^2 \end{aligned}$$

This result is valid for $h\nu \ll mc^2 = 511 \text{ keV}$, and an electron at rest or moving with $v \ll c$.

3 Compton scattering

For high-energy photons it is important to include the photon momentum and regard the scattering process as a particle collision. Then we have:

Energy conservation:

$$mc^2 + h\nu = \gamma mc^2 + h\nu' \quad (2)$$

and momentum conservation: x -component

$$\frac{h\nu}{c} = \gamma mv \cos \phi + \frac{h\nu'}{c} \cos \theta \quad (3)$$

and y -component

$$0 = \gamma mv \sin \phi - \frac{h\nu'}{c} \sin \theta \quad (4)$$

Now square equations (3) and (4) and add:

$$\begin{aligned} (\gamma mv \cos \phi)^2 + (\gamma mv \sin \phi)^2 &= (\gamma mv)^2 = \left(\frac{h\nu'}{c} \sin \theta\right)^2 + \left(\frac{h\nu'}{c} \cos \theta - \frac{h\nu}{c}\right)^2 \\ &= \left(\frac{h\nu'}{c}\right)^2 + \left(\frac{h\nu}{c}\right)^2 - \frac{2h^2\nu\nu'}{c^2} \cos \theta \end{aligned}$$

But

$$\beta^2 = 1 - \frac{1}{\gamma^2} \Rightarrow \gamma^2 \beta^2 = \gamma^2 - 1$$

So

$$(\gamma^2 - 1) \left(\frac{mc^2}{h\nu}\right)^2 = 1 + \left(\frac{\nu'}{\nu}\right)^2 - 2\frac{\nu'}{\nu} \cos \theta \quad (5)$$

Then squaring equation (2) gives:

$$(\gamma - 1)^2 \left(\frac{mc^2}{h\nu}\right)^2 = \left(1 - \frac{\nu'}{\nu}\right)^2 = 1 + \left(\frac{\nu'}{\nu}\right)^2 - 2\frac{\nu'}{\nu} \quad (6)$$

Subtracting equations (6) and (5), we get:

$$2(\gamma - 1) \left(\frac{mc^2}{h\nu}\right)^2 = 2\frac{\nu'}{\nu} (1 - \cos \theta)$$

Then using equation (2) again gives:

$$\begin{aligned} \frac{\nu'}{\nu} (1 - \cos \theta) &= \left(\frac{mc^2}{h\nu}\right)^2 \frac{h\nu}{mc^2} \left(1 - \frac{\nu'}{\nu}\right) \\ \frac{\nu'}{\nu} \left(1 - \cos \theta + \frac{mc^2}{h\nu}\right) &= \frac{mc^2}{h\nu} \end{aligned}$$

and so

$$\begin{aligned}\nu' &= \nu \left(\frac{mc^2}{h\nu} \right) \frac{1}{1 - \cos \theta + \left(\frac{mc^2}{h\nu} \right)} \\ &= \frac{\nu}{1 + \frac{h\nu}{mc^2} (1 - \cos \theta)}\end{aligned}$$

The differential cross section is modified by the frequency shift:

$$\left. \frac{d\sigma}{d} \right|_{QM} = r_0^2 (\hat{\mathbf{e}}_0 \cdot \hat{\mathbf{e}}_i^*)^2 \left(\frac{\nu'}{\nu} \right)^2$$